GEL105: IGNEOUS PETROLOGY

Disney Pixar: LAVA

2016

Instructor: Sarah Lambart

- Office hour: Thursday from 12 to 1pm + by appointments in my office (room 1205A)
- TA: Kevin Schrecengost and Allie Rubin
 - Office hour: in room 1362
 - Kevin: Tu/Th 10-11am; Allie: Tu/Th 9-10am
- Class website: Smartsite

www.smartsite.edu

 Textbook: Winter (Principle of Igneous and Metamophic petrology)

Labs:

- Twice a week for 3 hours each session
- 7 labs largely based on California rocks
- Term project (will be introduced in two weeks)
- Grading: 4 units ~60% lecture/40% labs
 - 2 midterms
 - Final exam
 - 5 problem sets (on lecture material)
 - Lab/lab quizzes
 - Term project
 - Figure presentation
- Late work: 25% per late day (for PS only!)

What is Igneous Petrology?









Plate tectonic vs. Volcanism



Note: this does not include mid-ocean ridge volcanism

Smithsonian Global Volcanism project - http://www.volcano.si.edu/world/find_regions.cfm



Crustal Plate Boundaries





How do you melt rocks?

 Heat them (e.g., ice at room temperature)
Decompress them (lower pressure)
Change the composition (e.g., add salt to ice / water to rock)



How do you create igneous rocks?

 Heat them (e.g., ice at room temperature)
Decompress them (lower pressure)
Change the composition (e.g., add salt to ice / water to rock)



Temperature decrease ⇒ crystallization

Goals of this class:

 Learn how to use the textural and chemical information of rocks to understand origin and process

Relate this to the bigger picture

LECTURE 1: EARTH'S INTERIOR

<4.44 billion years ago

Hydrodynamic gas loss

Maselve early atmosphere Hydro Degassing

> Impact degassing

Atmospheremagma ocean equilibration

Undegassed material?

gina ocean

Core

Canospheric mantle

Modern atmosphere

Continental crust

Convecting mantle

> Mid-ocean ridge

> > Ocean
> > island

Today

Earth's interior:

Chemical

Mechanical



Figure 1.3 in Winter



Crust

Continental crust



Oceanic crust



sediments (red clays) pillow lavas (basalt) sheeted dike complex gabbros

Moho discontinuity

Mantle (Peridotite)

Mechanical layers

Geothermal gradient: non linear with depth

Depth



Figure 1.9 in Winter

Mechanical layers



LET'S BUILD A PLANET...

Earth's origin

- How did Earth form?
- How did we get the layered planet we have today?
- What is the bulk composition of Earth?
- When did all of this happen (and how long did it take)?

Orion Nebula • M42

HST - ACS/WFC

Few dates

- Universe origin: Big bang – 12 to 15 Ga
- Milky way: 10 Ga
- Solar system: 4.56 Ga (from solar nebula): 99.9% of the mass collapses to form the Sun
 - 4.54 Ga: Earth's formation





Few dates

Mars-size impactor: Moon formation





Magma ocean



Magma ocean: 1st magma chamber



© 2000 Don Dixon / cosmographica.com

Why a layered Earth?

Goldschmidt classification (1937)



To first order, distribution between core and mantle: equilibrium partitioning between metal liquid and silicates (but no separate sulfide phase)

Why a layered Earth?

• Simple model





Fe, Mg, Si and O: account for 93% of the mass of Earth

How do we know all this?



Fig. 9. Meteorite collecting in Antarctica, during the 1978-79 season. Here the specimen is a tiny fragment of carbonaceous chondrite. Kazuyuki Suraishi of the National Institute of Polar Research, Tokyo, photographs the meteorite before moving it; Ursula Marvin of the Harvard-Smithsonian Center for Astrophysics holds a scale and numbering device that will appear in the photograph for identification. This photograph is by W. A. Cassidy.

Carbonaceous Chondrites = Bulk Earth Composition



BSE image of a chondrule Credit: Steven Simon

AND NOW?..

Sources of energy

Remnant heat from Earth's accretion

 Radioactive decay



Works



NEXT TIME

1001 recipes to create an igneous rocks...

TO READ: Chapter 1 (recap of lecture 1) Chapter 2 (help for Lab 1)